



... for a brighter future

Supernovae Simulations and Strategies: Application to the Dark Energy Survey

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*214th Meeting of the
American Astronomical Society
Pasadena, CA 2009-06-10*

Outline

- DES-Supernovae (DES-SN) Overview
- SNANA: SN light curve simulation & fitter
- DES-SN Selection Bias
- DES-SN Non-Ia Contamination
- SNANA in the Infrared
- Summary & Conclusions



DARK ENERGY
SURVEY

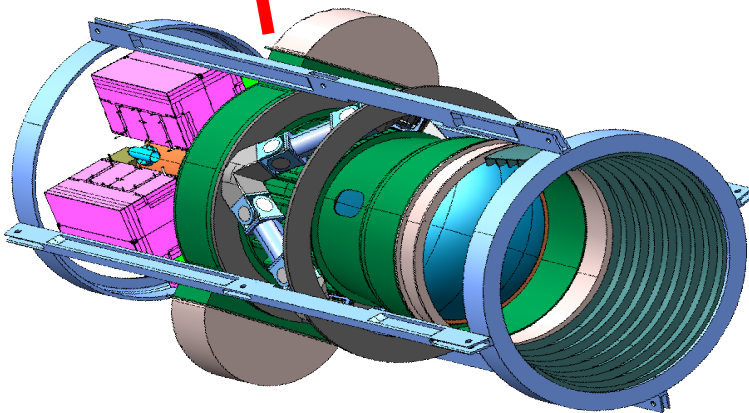
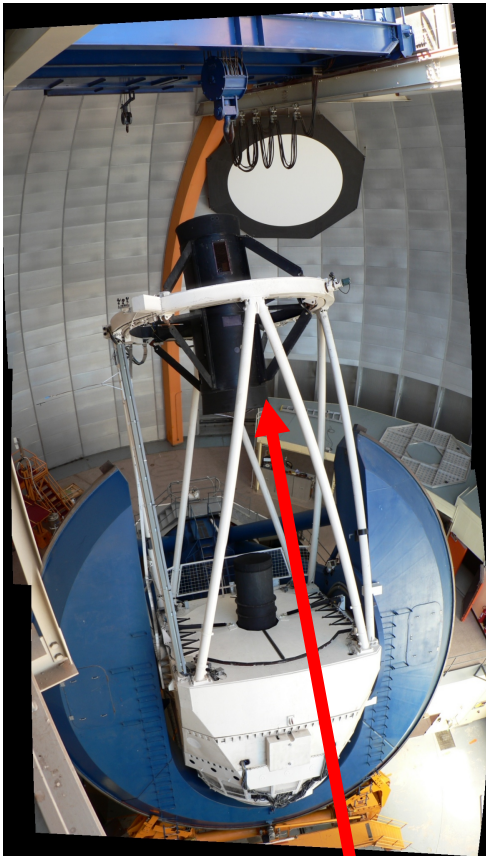


University of Chicago

Dark Energy Survey (DES)

DES is providing a new 520Mpixel CCD camera (DECam) for the Blanco 4m telescope in Chile in exchange for 525 survey nights over 5 year period for a 5000 square degree survey starting in 2011

DES uses thicker CCDs from Lawrence Berkeley National Laboratory with increased red sensitivity



DES Supernovae

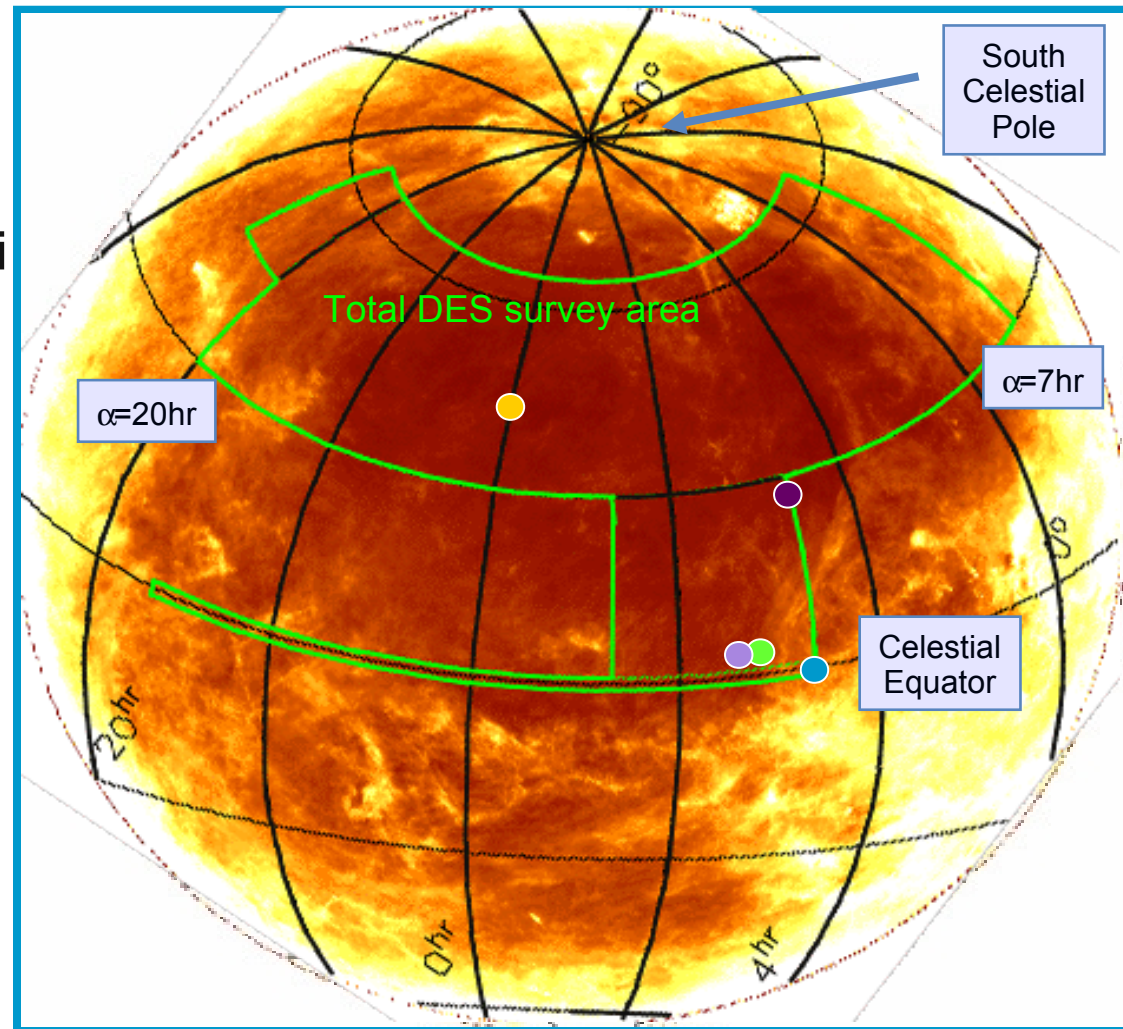
- DES time allocation fixes total supernovae (SNe) exposure time
 - 1260 hr planned over 5-year survey
 - maximal use of non-photometric time (~920 hr or 73%) planned
- Time per field & number of fields can be simulation optimized
 - ultra-deep strategy (3 square degrees = 1 DES field)
 - deep strategy (9 square deg.)*
 - shallow but wide strategy (27 square deg.)
 - hybrid strategy, e.g., 2 deep + 3 wide (15 square deg.)
- Results show hybrid strategy is best (more later)

* Highlighted in DES DOE proposal

Currently Favored DES-SN Fields

- Chosen to maximize:
 - visibility from DES site
 - past observation history
 - visibility from, e.g., Hawaii

Chandra Deep Field – South ●
 Sloan Stripe 82 ●
 SN Legacy Survey (SNLS) D1 ●
 XMM-Newton LSS ●
 ELAIS S1 ●



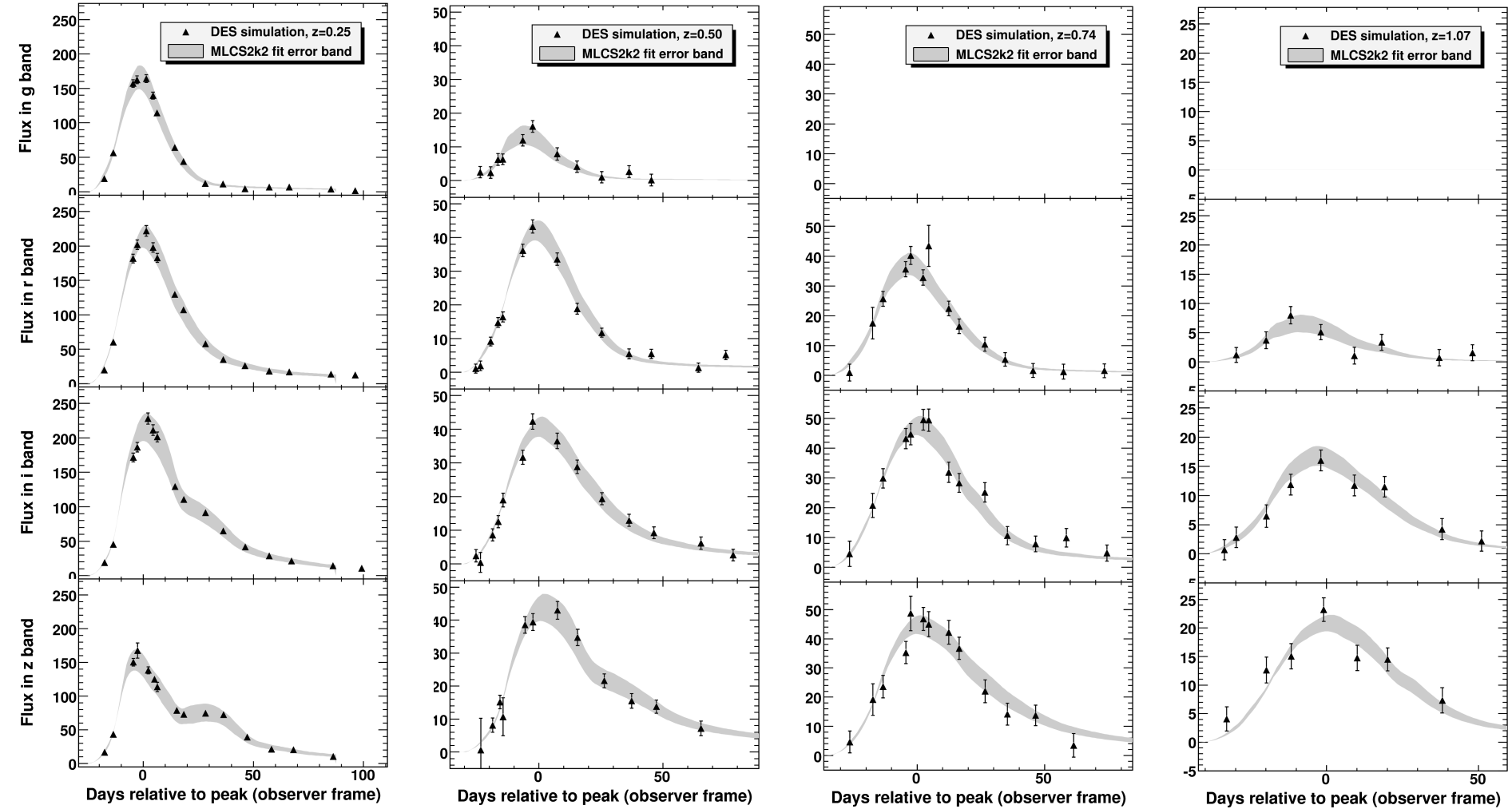
From a study by Peter Nugent

SNANA: SuperNova ANAlysis package for DES

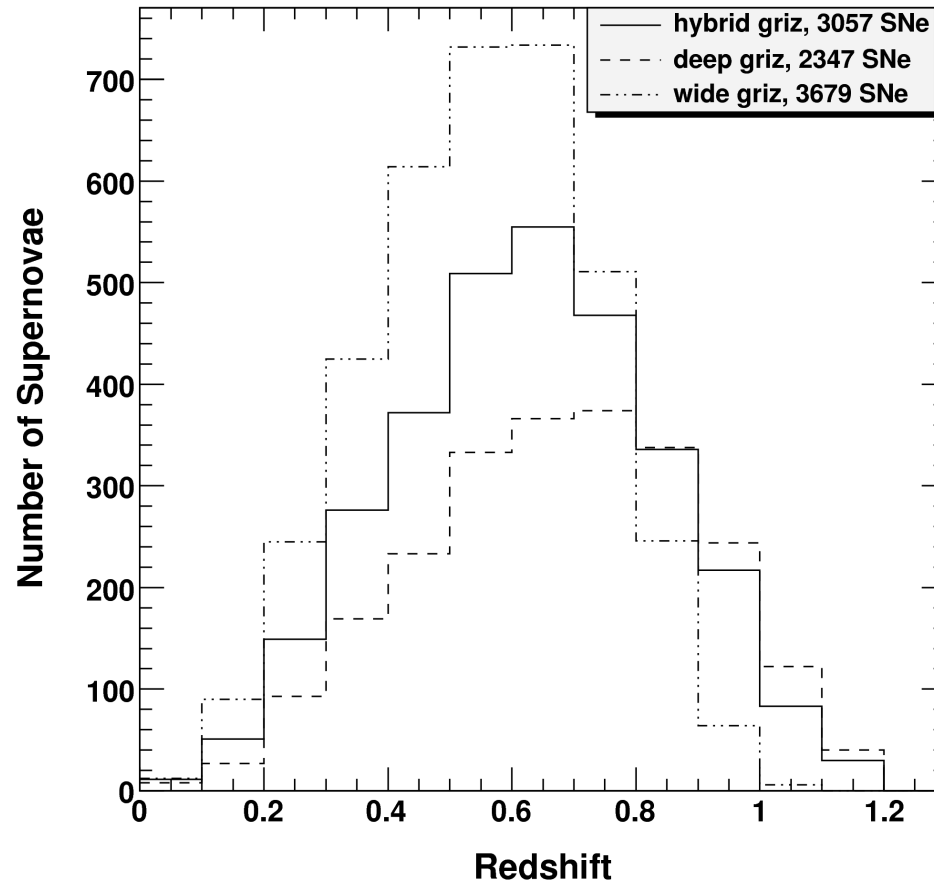
R. Kessler (U. Chicago), J. P. Bernstein, S. Kuhlmann, & H. Spinka (ANL)

- Public URL: <http://www.sdss.org/supernova/SNANA.html>
- Also used by SDSS & LSST
- Software suite for simulating and fitting SN light curves
- Motivation was a more accurate and complete study of DES-SN capabilities including DES CCD and filter characteristics, CTIO sky fluctuations using Essence data inputs, dust extinction effects, etc.
- Uses various models (e.g., MLCS2k2, SALT-II, stretch, etc.)
- Models and fits both Ia and non-Ia SNe

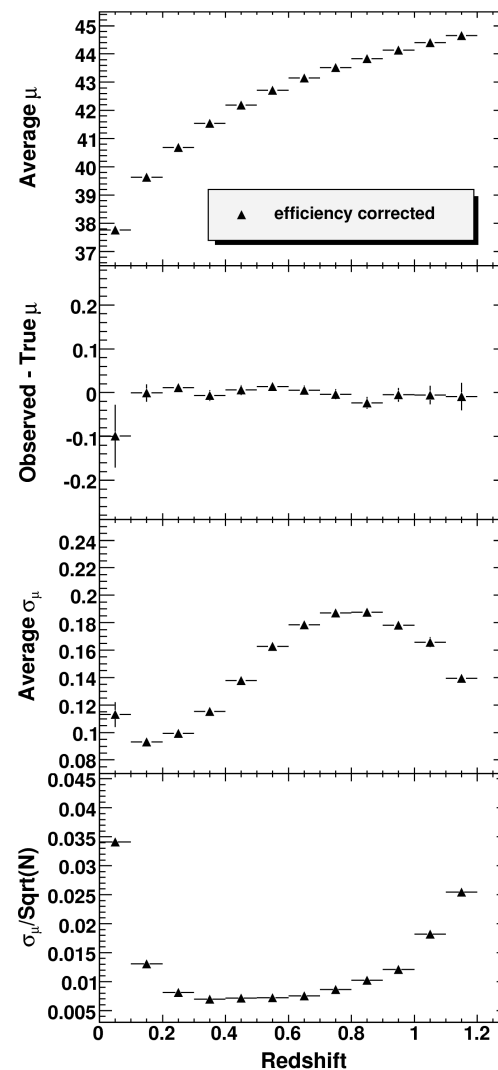
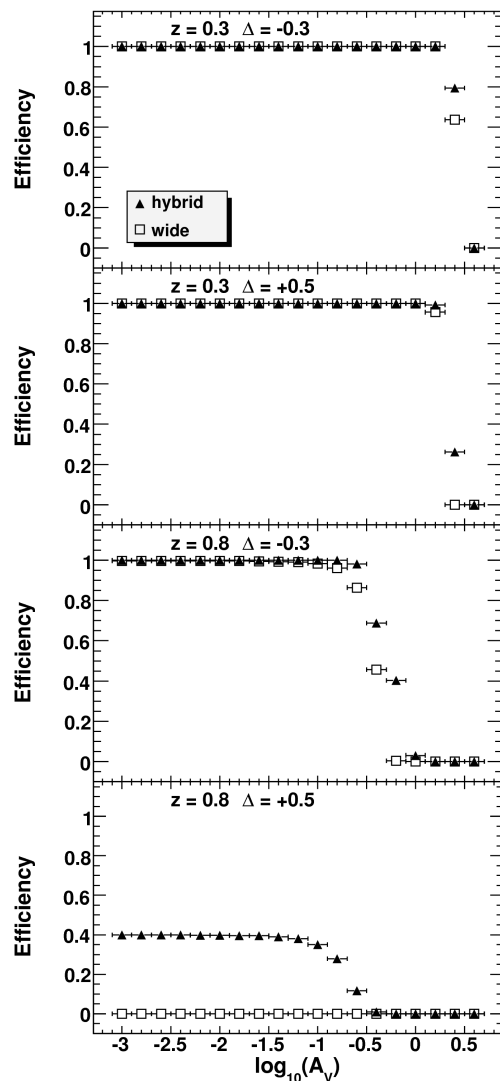
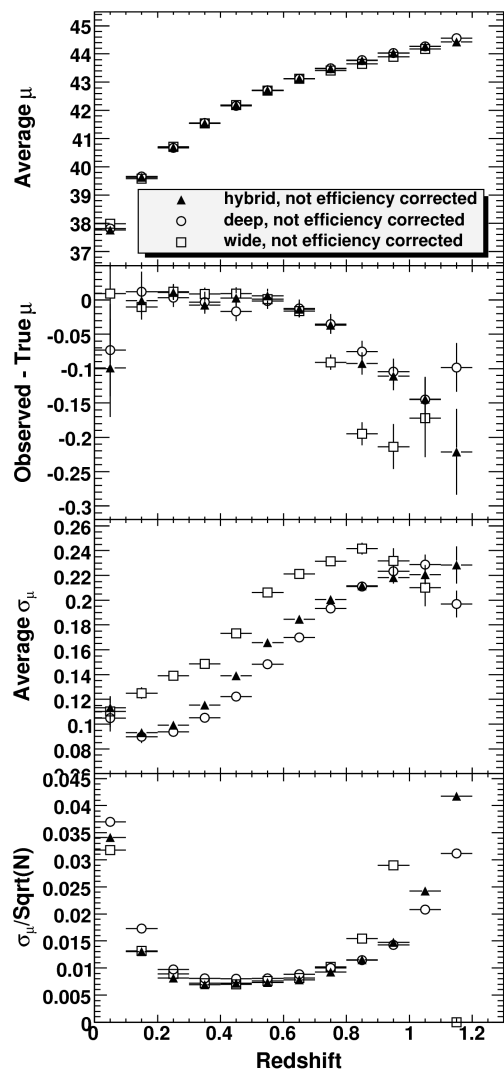
Example Simulated MLCS2k2 DES SN Ia Light Curves



Number Of Supernovae For Different Strategies

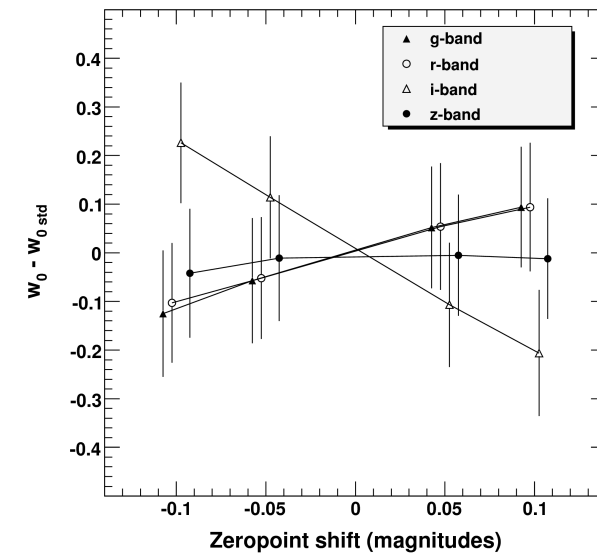
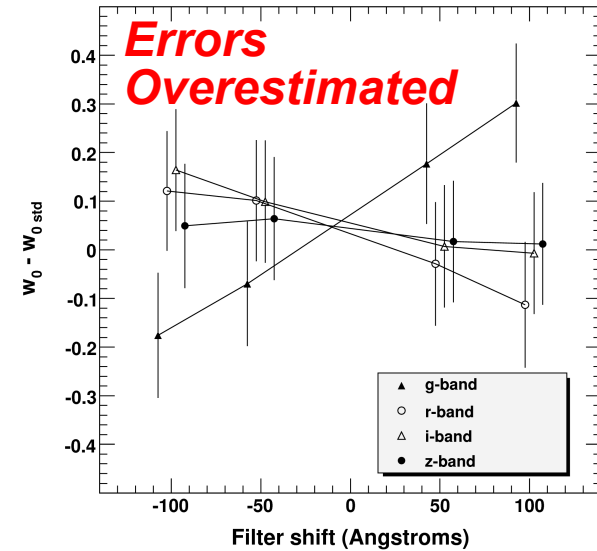
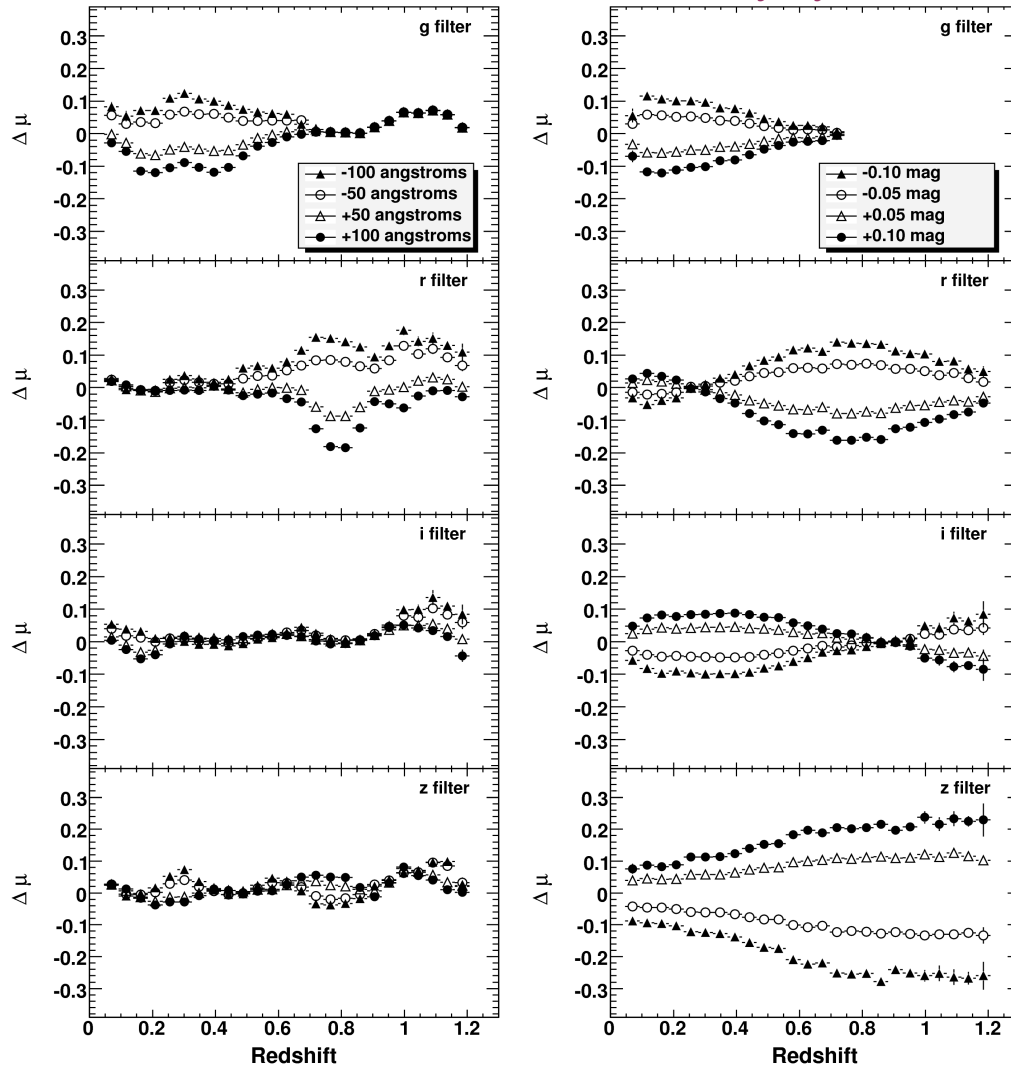


Selection Bias And The Hubble Diagram Projection

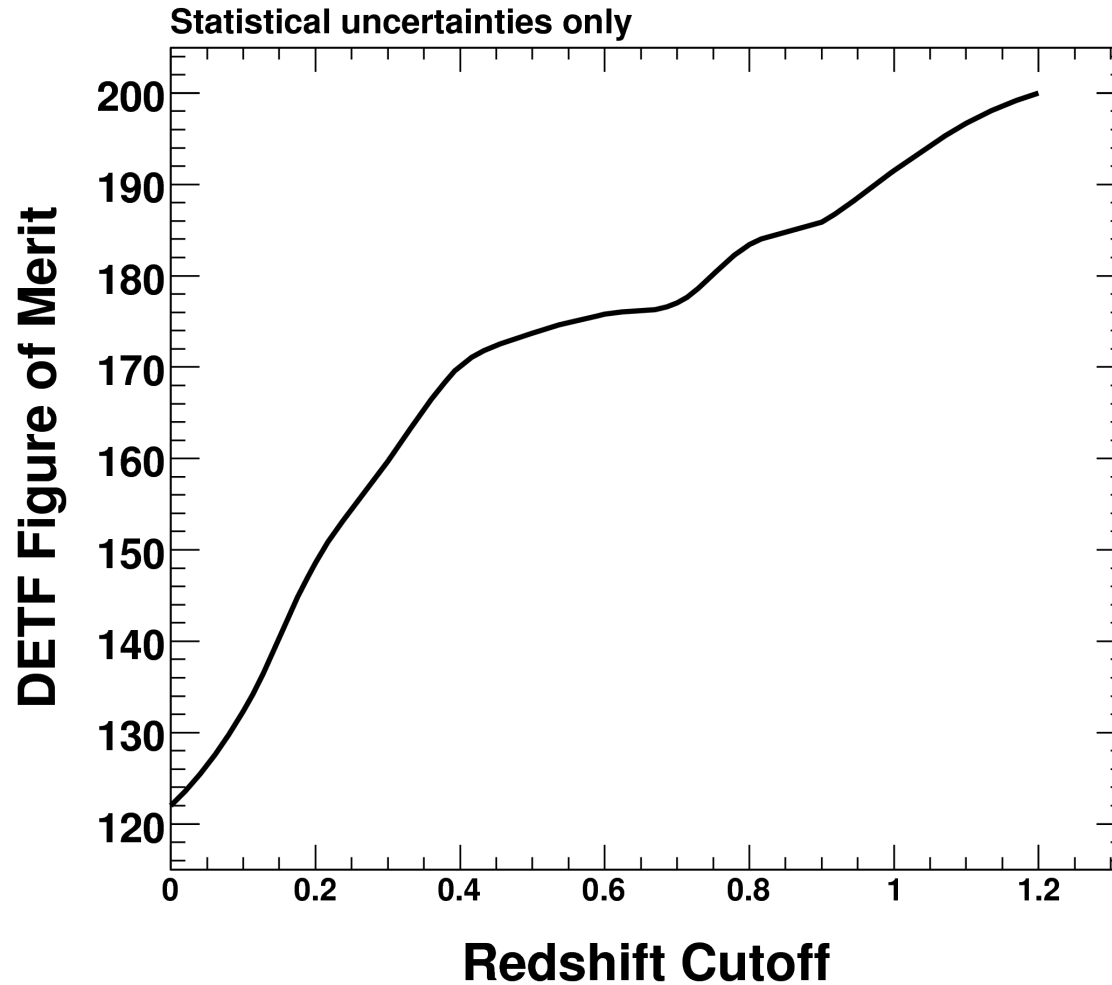


Effect of Filter Centroid and Zeropoint Shifts

Study by Ribamar Reis



Spectroscopic Redshift Cutoff and DETF Figure of Merit



A COMPARATIVE STUDY OF THE ABSOLUTE MAGNITUDE DISTRIBUTIONS OF SUPERNOVAE

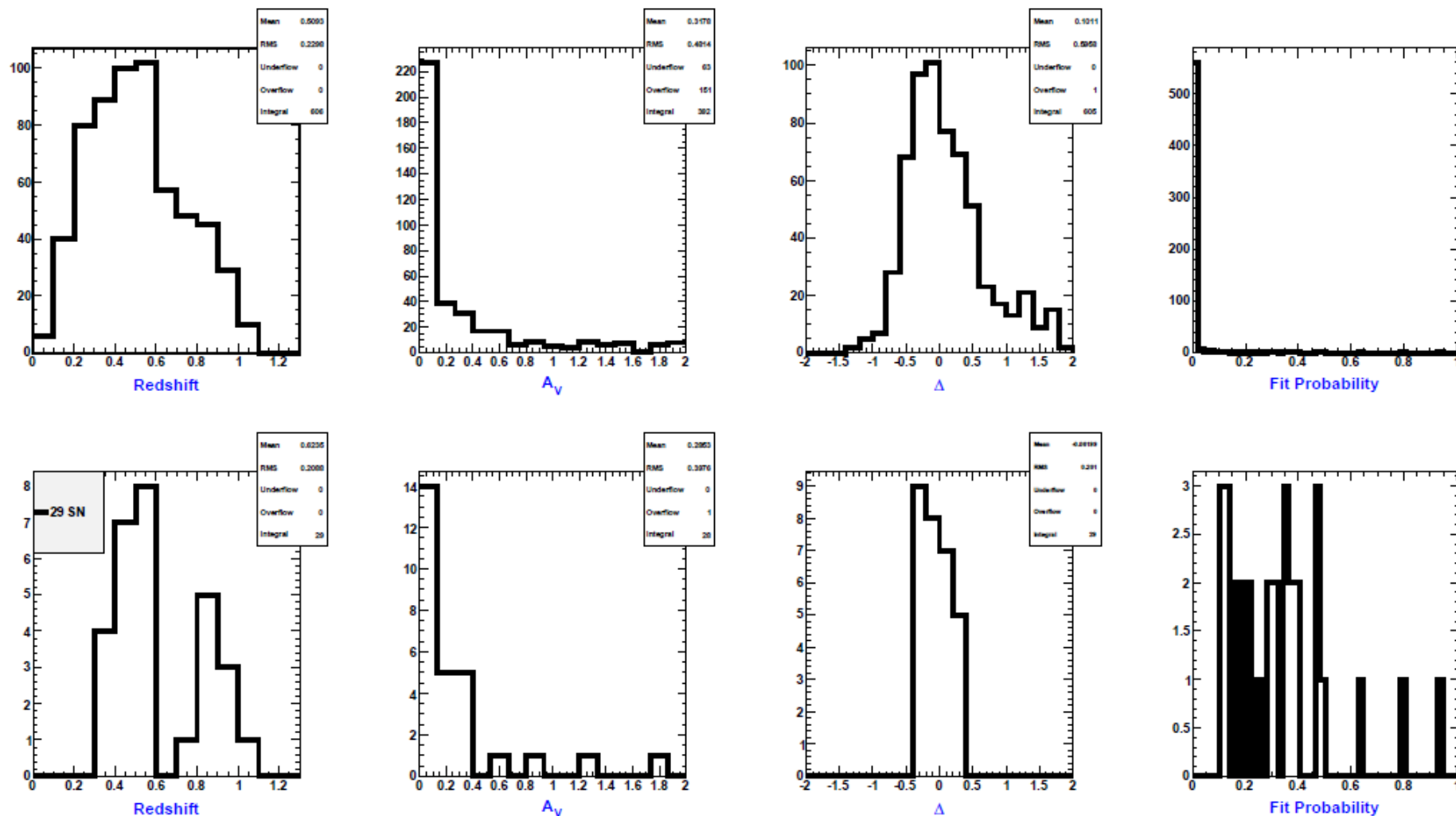
DEAN RICHARDSON, DAVID BRANCH, DARRIN CASEBEER, JENNIFER MILLARD, R. C. THOMAS, AND E. BARON
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Received 2001 September 24; accepted 2001 October 24

TABLE 1
RESULTS

SN Type	$\overline{M}_{B,obs}$	σ_{obs}	$\overline{M}_{B,int}$	σ_{int}	Conf.	N
Normal Ia	-19.16 ± 0.07	0.76	-19.46	0.56	0.89	111
Total Ibc	-17.92 ± 0.30	1.29	-18.04	1.39	0.96	18
Bright Ibc.....	-19.72 ± 0.24	0.54	-20.26	0.33	~ 1	5
Normal Ibc	-17.23 ± 0.17	0.62	-17.61	0.74	~ 1	13
Total II-L.....	-17.80 ± 0.22	0.88	-18.03	0.90	0.91	16
Bright II-L	-19.12 ± 0.12	0.23	-19.27	0.51	~ 1	4
Normal II-L...	-17.36 ± 0.12	0.43	-17.56	0.38	~ 1	12
II-P	-16.61 ± 0.23	1.23	-17.00	1.12	~ 1	29
II _n	-18.78 ± 0.31	0.92	-19.15	0.92	~ 1	9

Fluctuate Nugent templates by numbers from Richardson et al.

Example: Nugent Type-Ibc Top 4 plots only SNR cuts



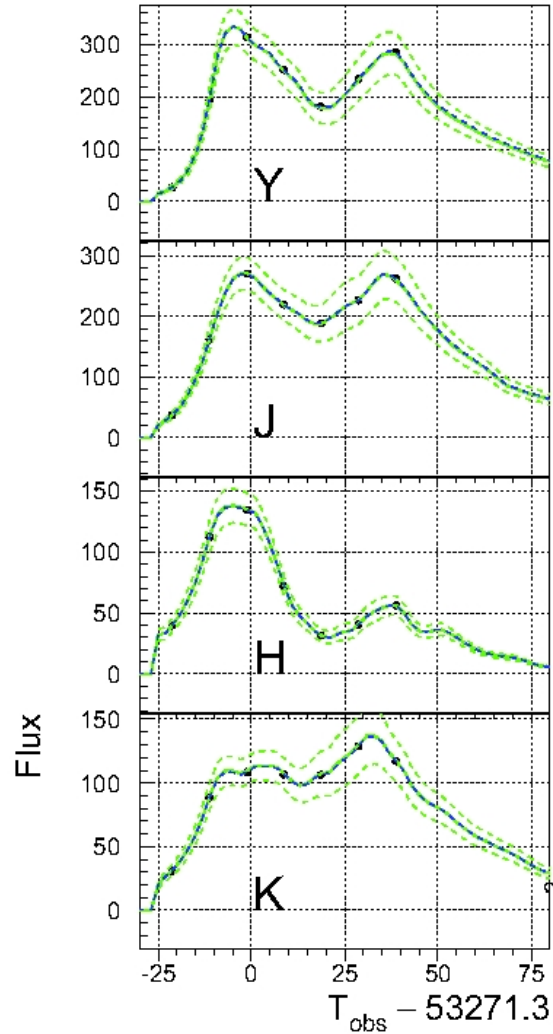
Bottom 4 plots include cut of fit probability > 0.1

SNANA IR Simulations

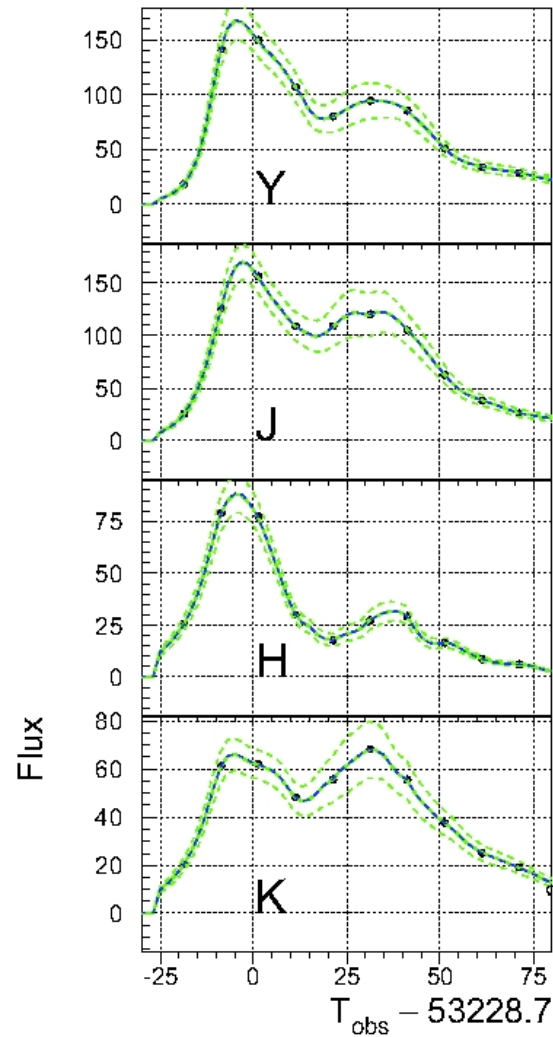
- New IR model for SNANA: mlcs2k2.IR
- UBVRIYJHK filters
- Uses new 9-filter model code
- UBVRI works as mlcs2k2.v006b if YJHK templates do not exist
- NB: A_V -prior dominates YJHK fits b/c sim has no lever arm on color

“Perfect” VIDEO IR SN Ia Light Curves

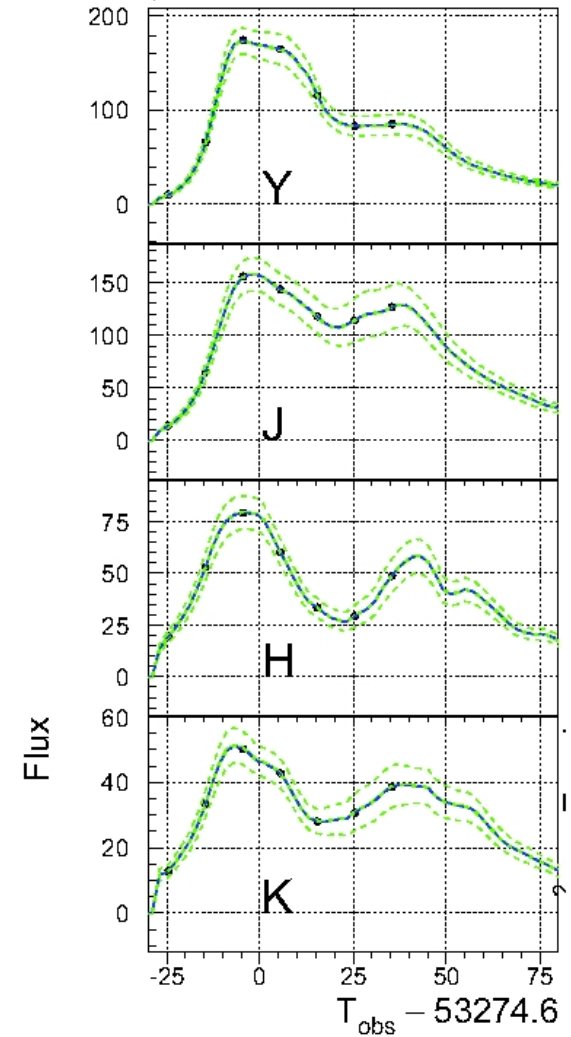
$z = 0.29$



$z = 0.34$



$z = 0.44$

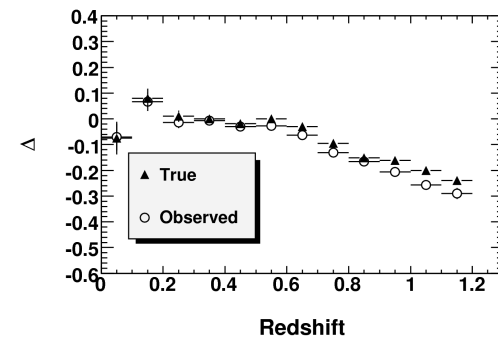
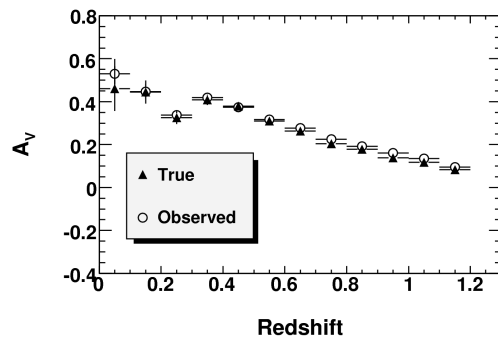
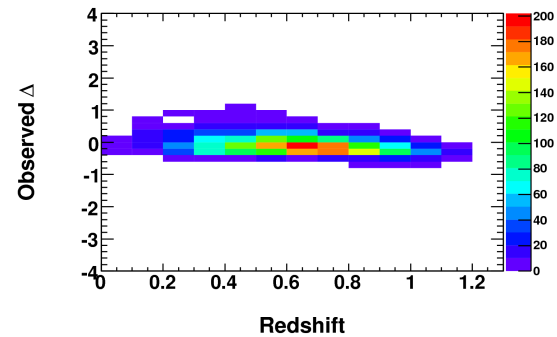
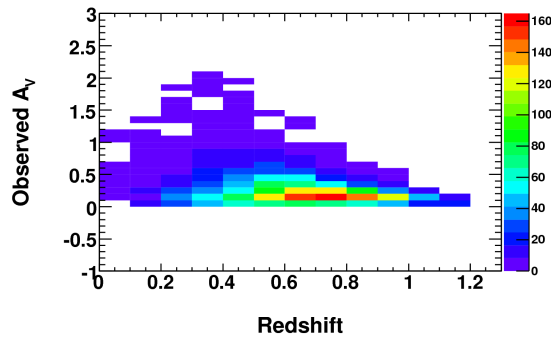
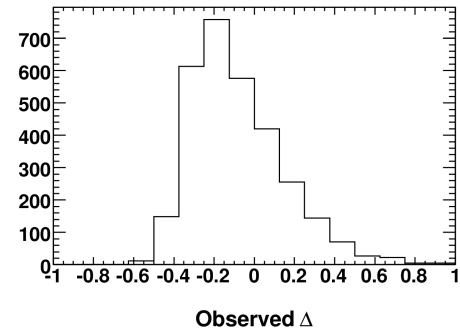
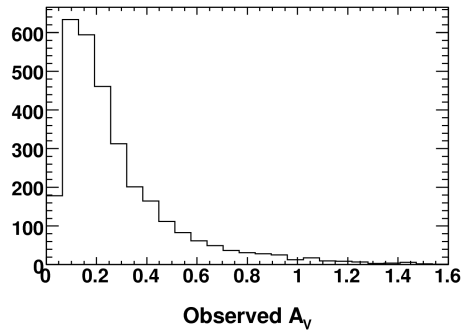


Summary & Conclusions

- DES will compile a sample of ~ 3000 well-measured SNe to $z \sim 1$
- Hybrid strategy of “deep” and “wide” fields optimal
- DES-SN Strategy simulation paper advancing
 - systematics, non-Ia, and IR studies will complete paper
 - journal submission planned this summer
- Initial SNANA IR model in place
- Follow-on DES/VIDEO IR SN paper next

Backup

Extinction (A_V) and Light Curve Shape Parameter (Δ)



Photometric Redshifts

